

EUMETSAT plans/status for EPS-SG microwave missions, CIMR and EPS Sterna

**Christophe Accadia, Jörg Ackermann,
Sabatino Di Michele, Imke Krizek,
Vinia Mattioli, Francesco De Angelis, Robin
Ekelund, Vasileios Barlakas, Linda Corucci**

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LEO Satellites: Retrieval and Applications 1-2 March 2023*



EPS-SG MW and sub-mm wave missions

MWS, MWI and ICI

CIMR at EUMETSAT

EUMETSAT activities in support of CIMR mission

EPS Sterna

Status and plans for the Sterna Constellation



EPS-SG: EUMETSAT Polar System – Second Generation

The EUMETSAT Polar System (EPS) in Low Earth Orbit (LEO) will be followed by a second generation system (EPS-SG). European contribution to the Joint Polar System set up with NOAA.

Same orbit as Metop (sun-synchronous, 832 km mean altitude, 09:30 local time of the descending node).

Payload distributed between the two parallel satellites Metop-SG A and B. Nominal lifetime of 7.5 years/spacecraft for an operational lifetime of the programme over 21 years.

▫

Metop payload	Metop-SG payload	Metop-SG satellite
Infrared Atmospheric Sounding Interferometer (IASI)	Infrared Atmospheric Sounding Interferometer – New Generation (IASI-NG)	A
Advanced Very High Resolution Radiometer (AVHRR)	Visible-Infrared Imager (METImage)	A
Advanced Microwave Sounding Unit A (AMSU-A1/A2), Microwave Humidity Sounder (MHS)	Micro-Wave Sounder (MWS)	A
Global Ozone Monitoring Experiment 2 (GOME-2)	UV-VIS-NIR-SWIR Sounder (Sentinel-5)	A
Advanced Scatterometer (ASCAT)	Scatterometer (SCA)	B
Global Navigation Satellite System Receiver for Atmospheric Sounding (GRAS)	Radio Occultation (RO)	A and B
-	Micro-Wave Imager (MWI)	B
-	sub-mm wave Ice Cloud Imager (ICI)	B
-	Multi-viewing, -channel, -polarisation Imager (3MI)	A

Metop-SG A



Metop-SG B

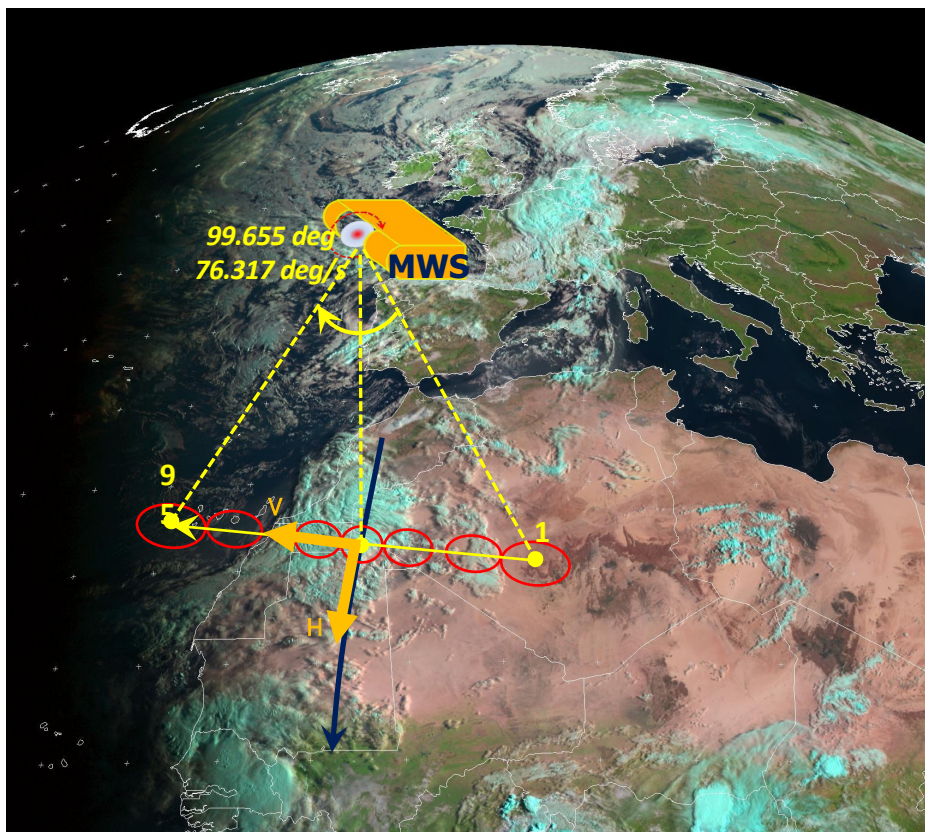




Microwave Sounding mission (MWS)

Test data version 2 now available at:

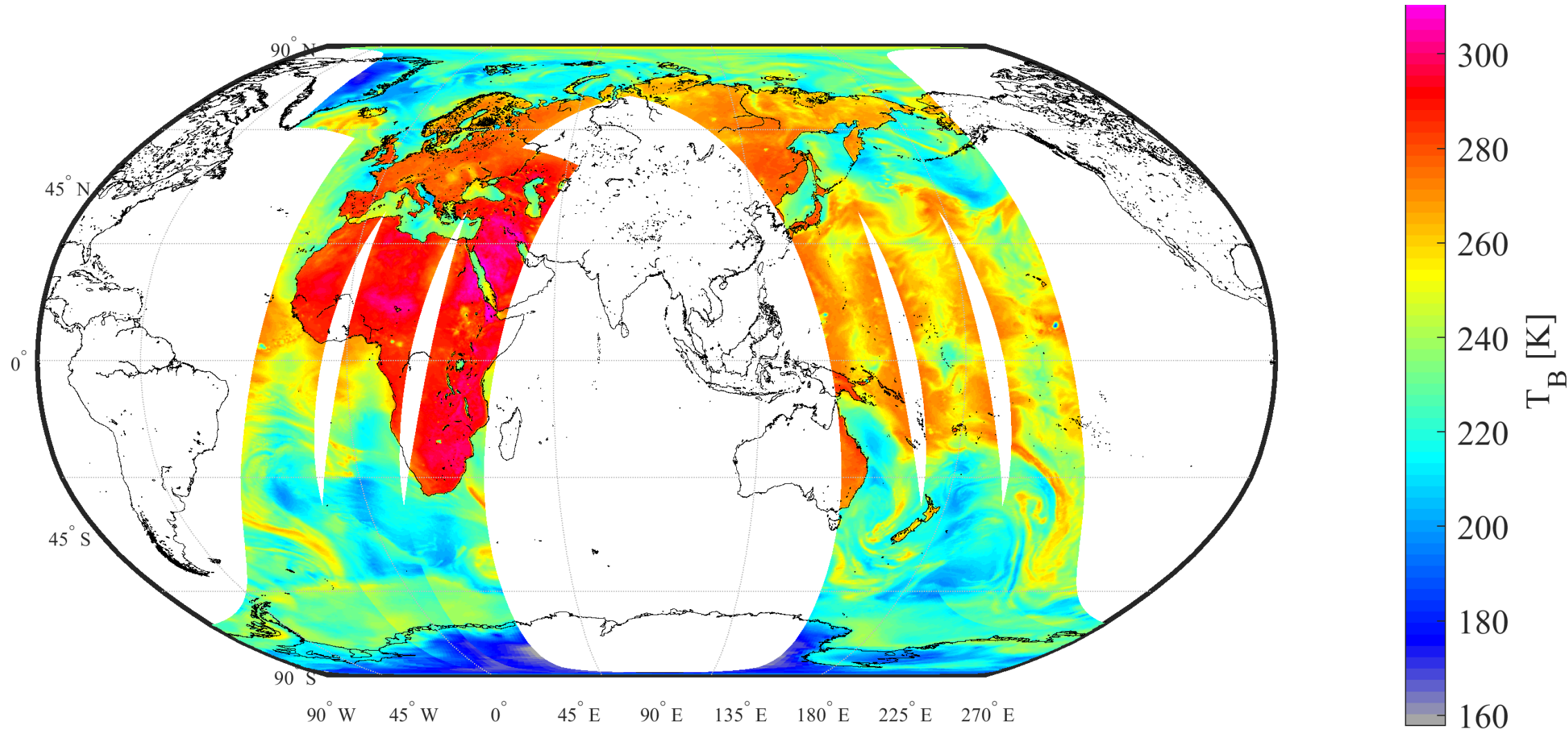
<https://www.eumetsat.int/new-version-eps-sg-mws-test-data-released>



Channel	Centre Frequency (GHz)	Bandwidth per passband (MHz)	Footprint (km)
MWS-1	23.8	270	40
MWS -2	31.4	180	40
MWS -3	50.3	180	40
MWS -4	52.8	400	40
MWS-5	53.246 ± 0.08	2x140	40
MWS -6	53.596±0.115	2x170	40
MWS-7	53.948 ± 0.081	2x142	40
MWS -8	54.40	400	40
MWS -9	54.94	400	40
MWS -10	55.50	330	20
MWS -11	57.290344	330	20
MWS-12	57.290344±0.217	2x78	20
MWS-13	57.290344 ±0.3222±0.048	4x36	20
MWS-14	57.290344±0.3222±0.022	4x16	20
MWS-15	57.290344±0.3222±0.010	4x8	20
MWS-16	57.290344±0.3222±0.0045	4x3	20
MWS-17	89.0	4000	17
MWS-18	164-167	3000	17
MWS-19	183.311±7.0	2x2000	17
MWS-20	183.311±4.5	2x2000	17
MWS-21	183.311±3.0	2x1000	17
MWS-22	183.311±1.8	2x1000	17
MWS-23	183.311±1.0	2x500	17
MWS-24	229	2000	17

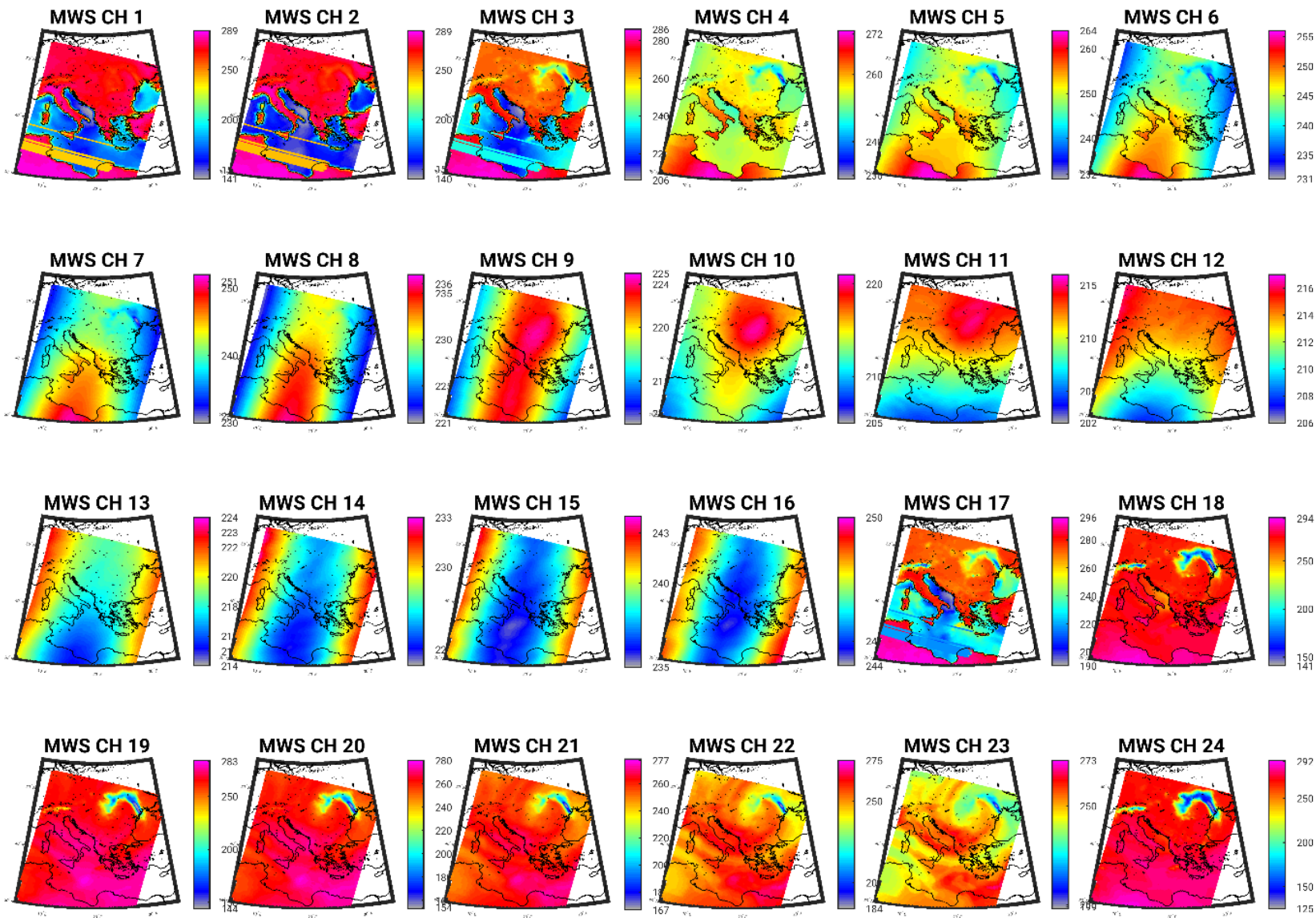


MWS Channel 17 (89 GHz) – Three consecutive orbits





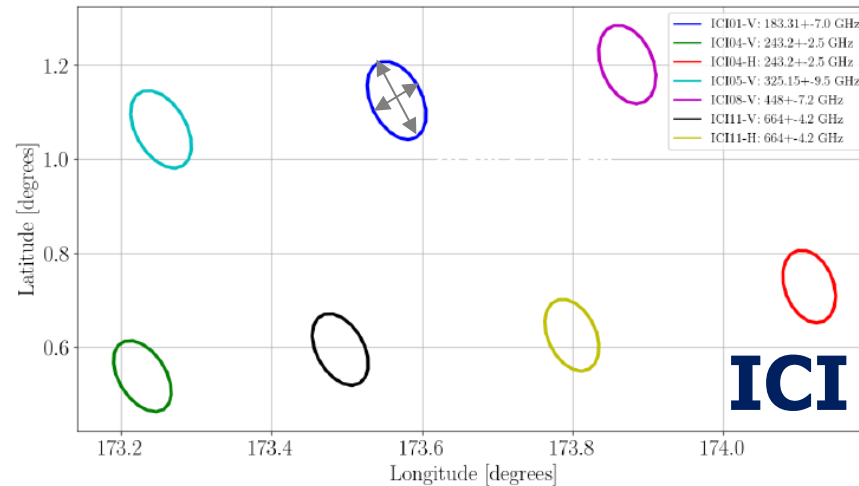
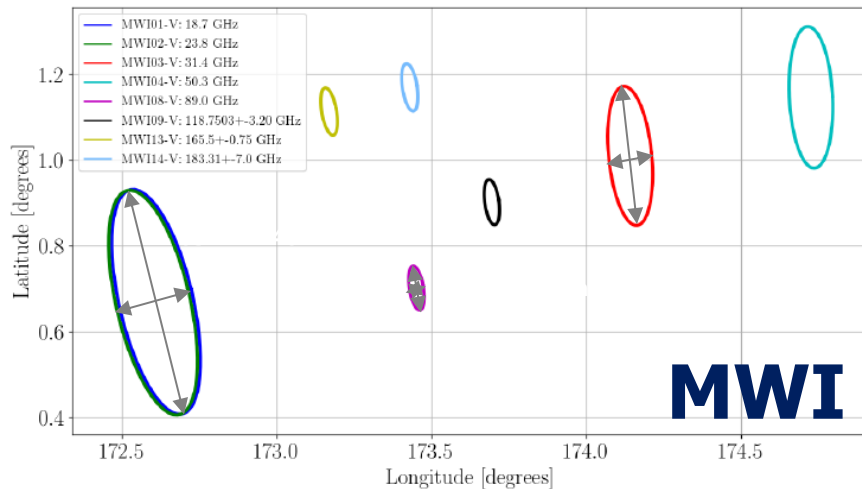
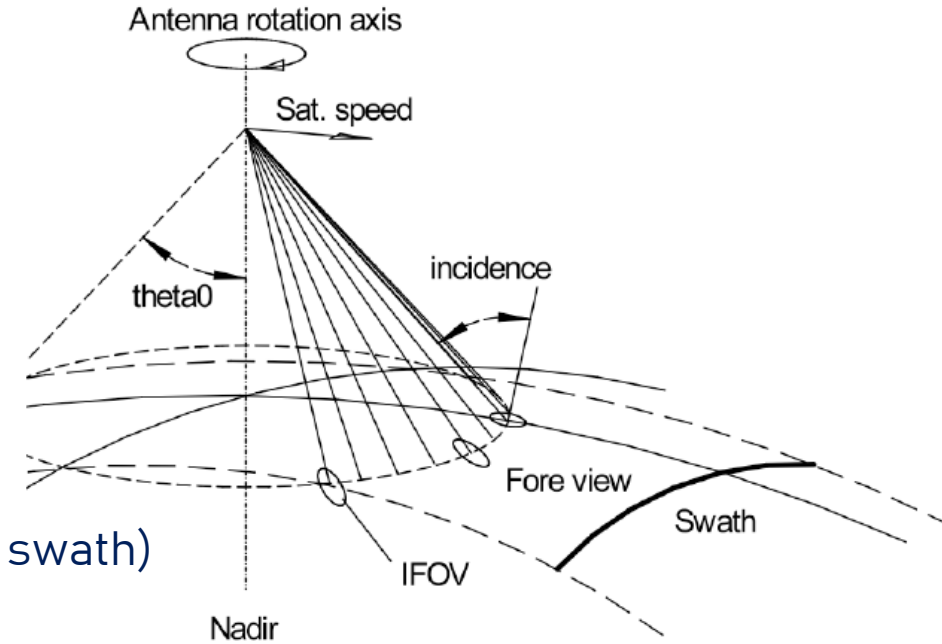
MWS L1B: TOA brightness temperatures





MW and Sub-mm Wave Imaging missions (MWI & ICI)

- ICI and MWI are conically scanning passive imagers
- Incidence angles within $53^\circ \pm 2^\circ$
- On the same platform, together with the scatterometer (SCA)
- Observations acquired $\pm 65^\circ$ in azimuth in the fore view (about 1700 km swath)





MWI and ICI in a nutshell

MWI

ICI

Channel	Frequency (GHz)	Bandwidth (MHz)	Polarisation	Footprint (km)	Channel	Frequency (GHz)	Bandwidth (MHz)	Polarisation	Footprint (km)
MWI-1	18.7	200	V, H	50	ICI-1	183.31±7.0	2x2000	V	16
MWI-2	23.8	400	V, H		ICI-2	183.31±3.4	2x1500	V	
MWI-3	31.4	200	V, H		ICI-3	183.31±2.0	2x1500	V	
MWI-4	50.3	400	V, H		ICI-4	243.2±2.5	2x3000	V, H	
MWI-5	52.61	400	V, H	ICI-5	325.15±9.5	2x3000	V		
MWI-6	53.24	400	V, H	ICI-6	325.15±3.5	2x2400	V		
MWI-7	53.75	400	V, H	ICI-7	325.15±1.5	2x1600	V		
MWI-8	89.0	4000	V, H	ICI-8	448±7.2	2x3000	V		
MWI-9	118.7503±3.2	2x500	V	ICI-9	448±3.0	2x2000	V		
MWI-10	118.7503±2.1	2x400	V	ICI-10	448±1.4	2x1200	V		
MWI-11	118.7503±1.4	2x400	V	ICI-11	664±4.2	2x5000	V, H		
MWI-12	118.7503±1.2	2x400	V						
MWI-13	165.5±0.75	2x1350	V	10					
MWI-14	183.31±7.0	2x2000	V						
MWI-15	183.31±6.1	2x1500	V						
MWI-16	183.31±4.9	2x1500	V						
MWI-17	183.31±3.4	2x1500	V						
MWI-18	183.31±2.0	2x1500	V						

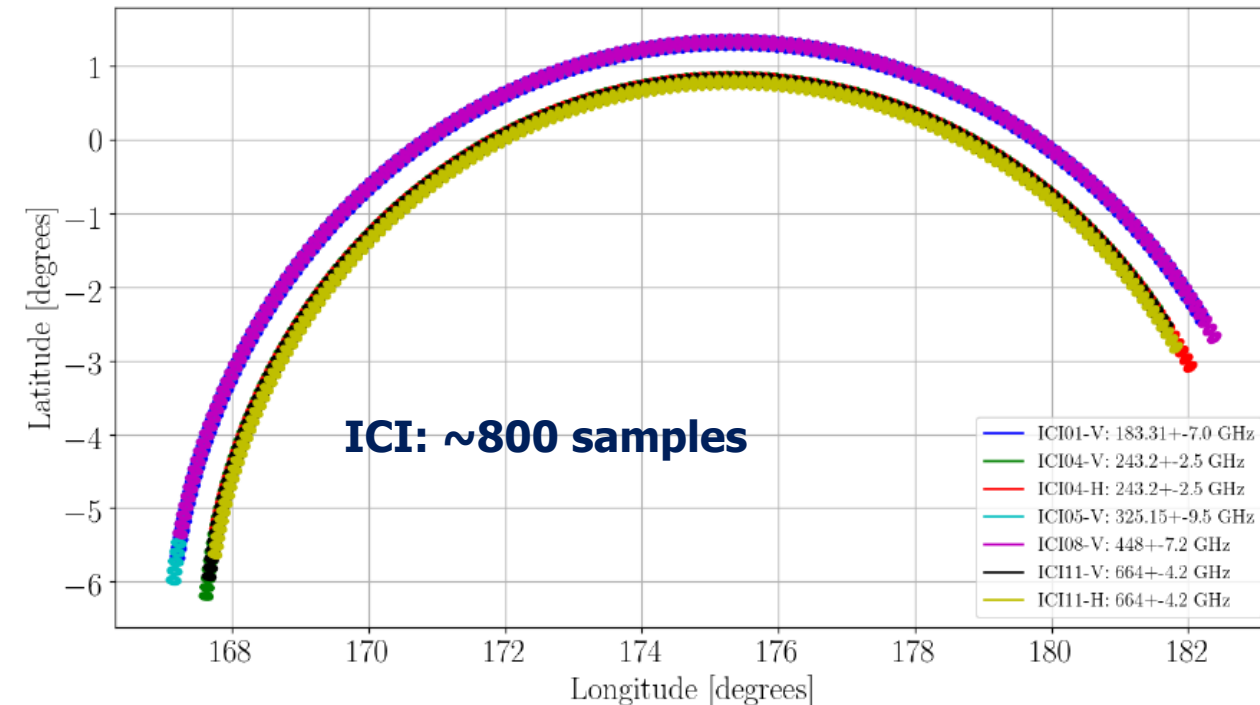
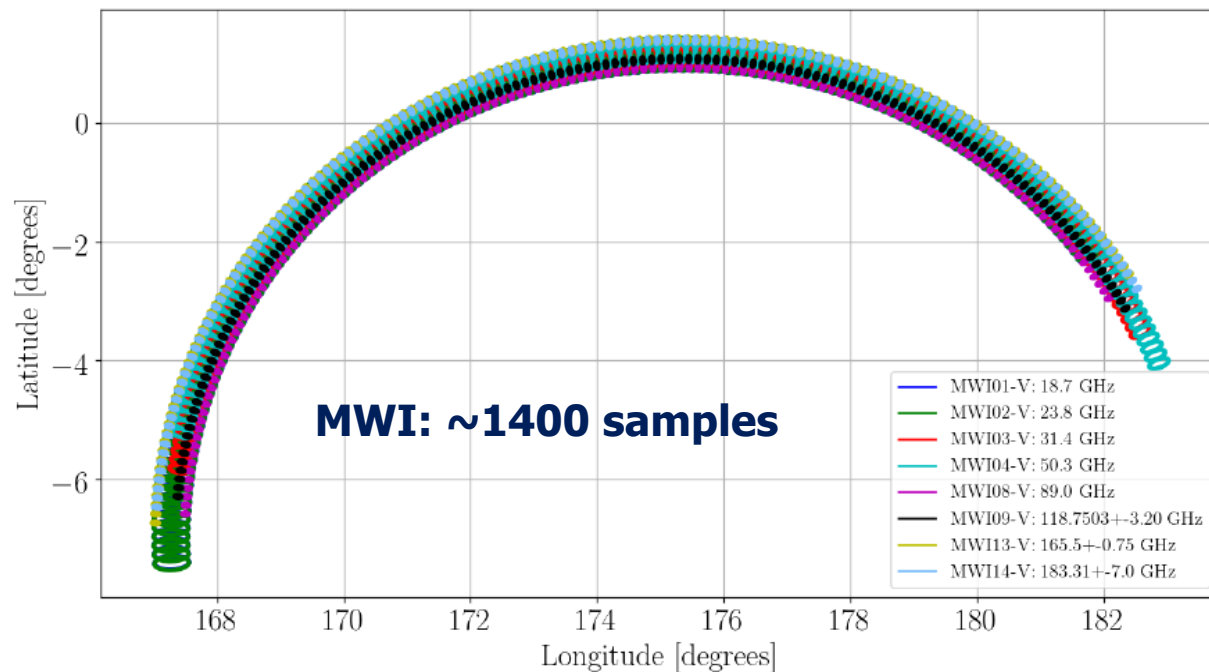


Innovative sampling of both imagers: new opportunities

For more details, test data version 2 now available at:

<https://www.eumetsat.int/new-versions-microwave-imager-and-ice-cloud-imager-l1b-test-data-released-nov-2022>

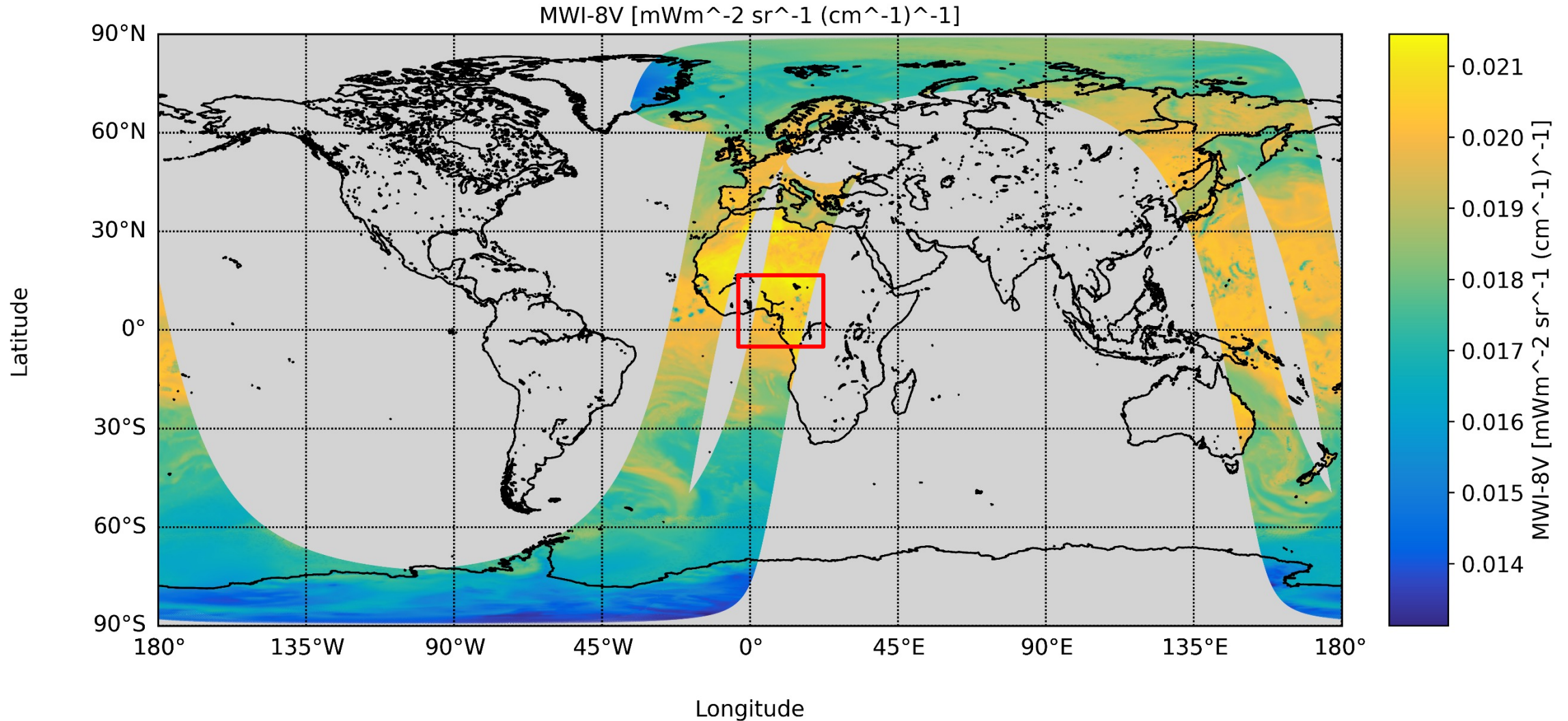
This is the pre-launch test dataset





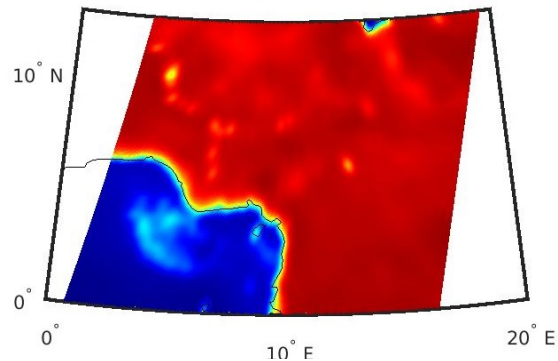
L1B TDP V2: MWI two summer consecutive orbits

MWI-8V

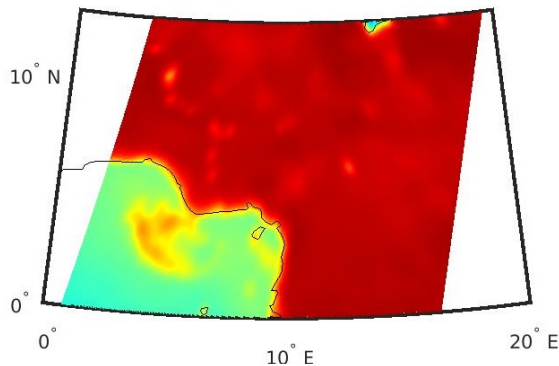


Overview of the MWI channels –T0A brightness temperatures

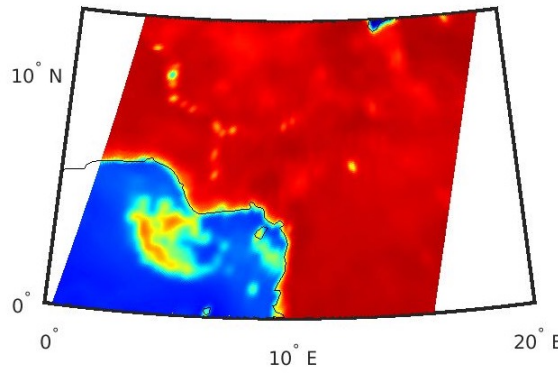
MWI-1H 18.7 GHz



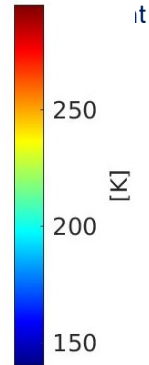
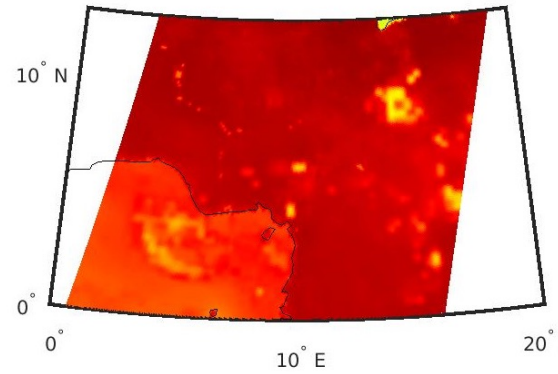
MWI-2H 23.8 GHz



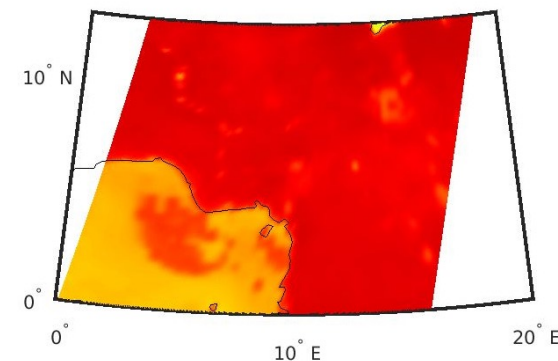
MWI-3H 31.4 GHz



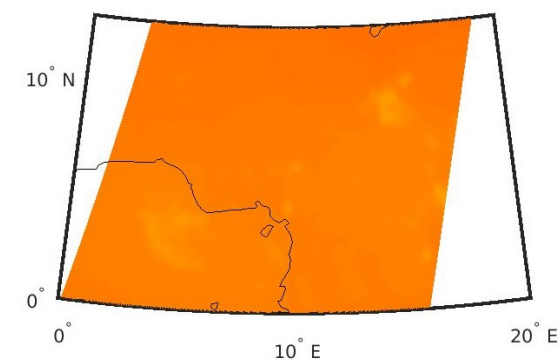
MWI-8H 89.0 GHz



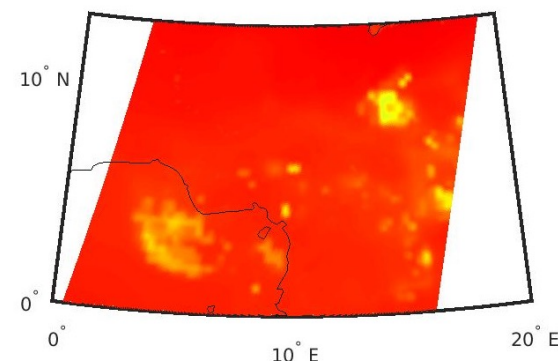
MWI-4H 50.3 GHz



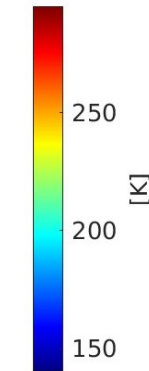
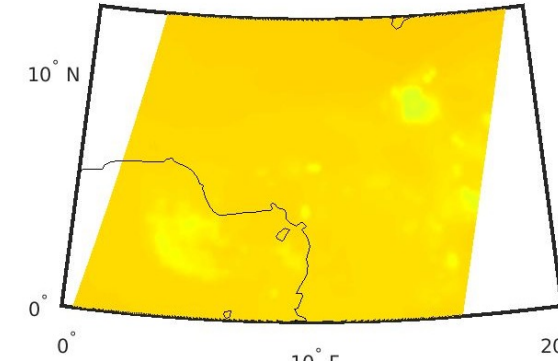
MWI-6V 53.24 GHz



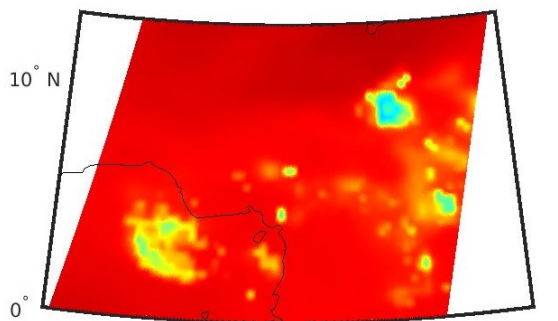
MWI-9V 118.75 GHz
 ± 3.20 GHz



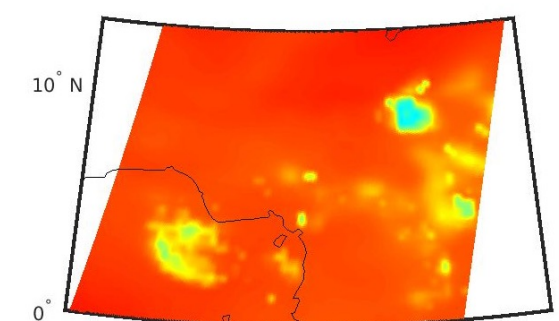
MWI-11V 118.75 GHz
 ± 1.40 GHz



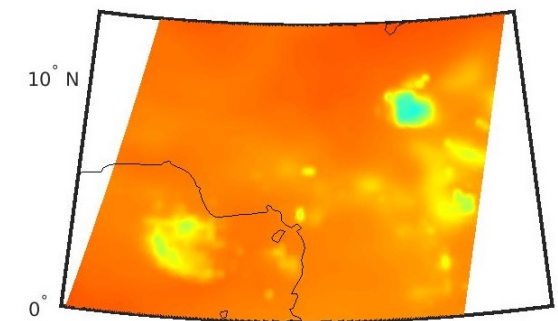
MWI-13V 165.5 GHz
 ± 0.75 GHz



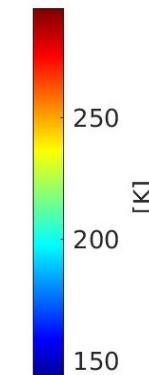
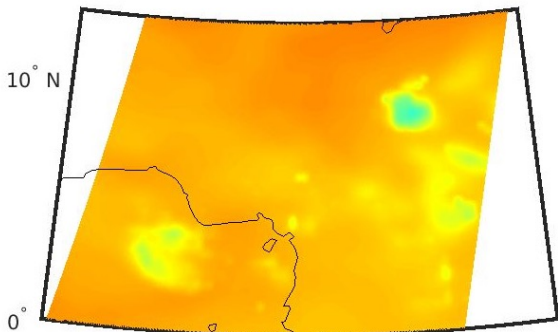
MWI-15V 183.31 GHz
 ± 6.10 GHz



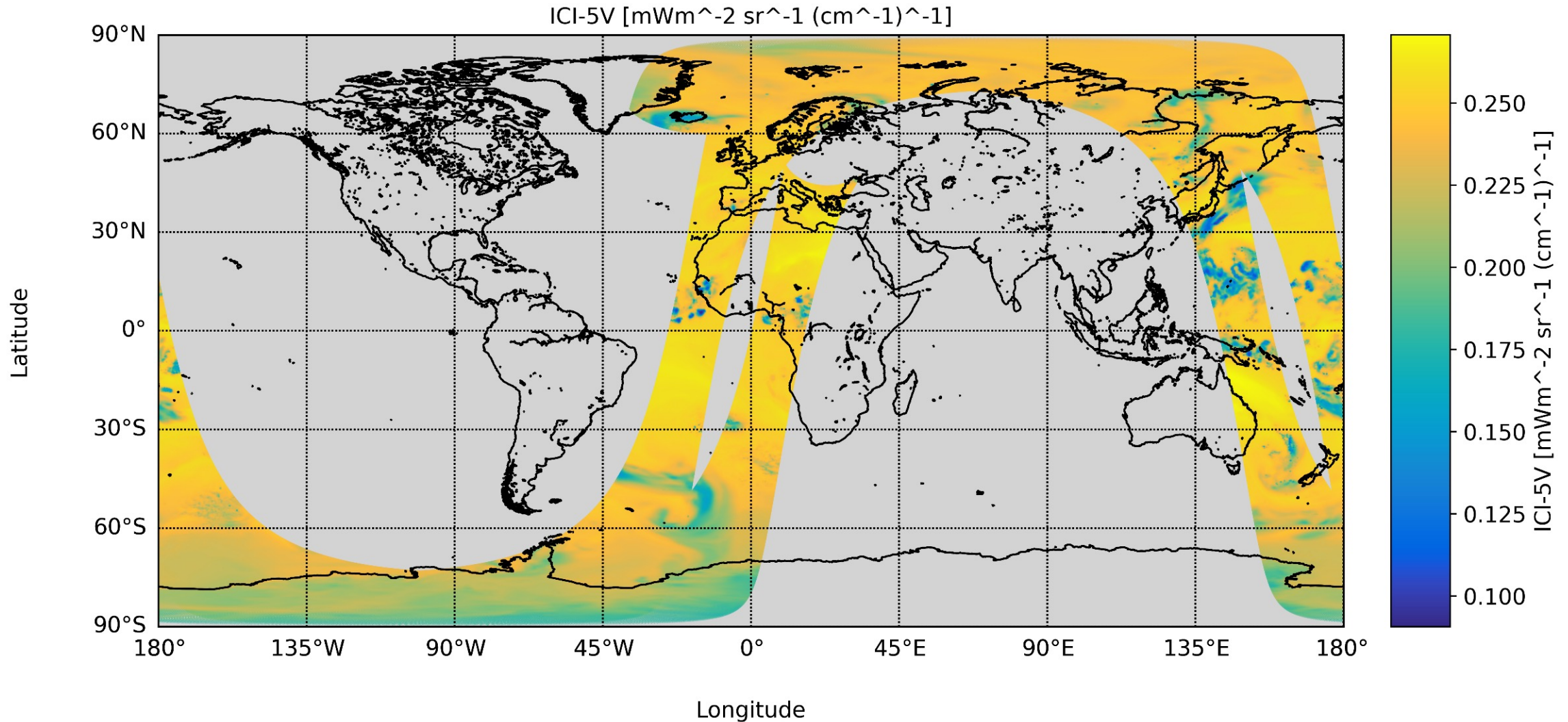
MWI-17V 183.31 GHz
 ± 3.40 GHz



MWI-18V 183.31 GHz
 ± 2.00 GHz



ICI-5V (325.15 ± 9.5 GHz)

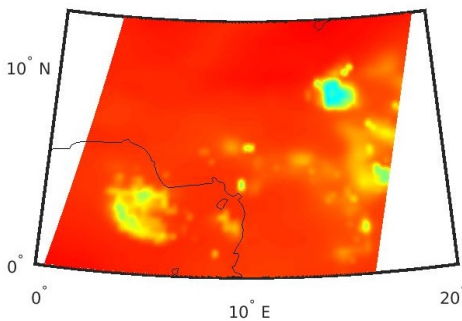




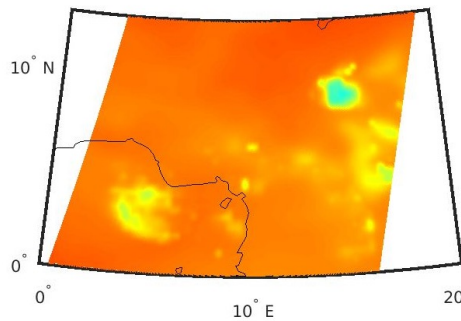
Overview of the ICI channels – TOA brightness temperatures

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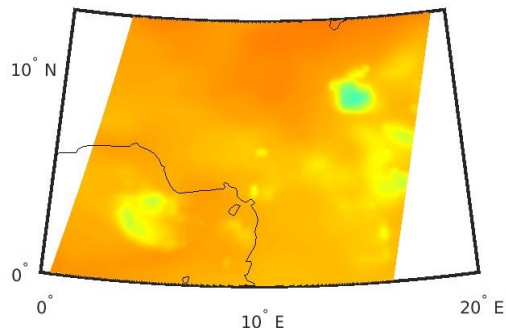
ICI-1V 183.31 GHz
 ± 7.00 GHz



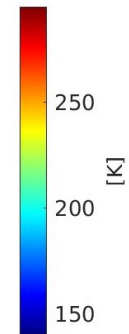
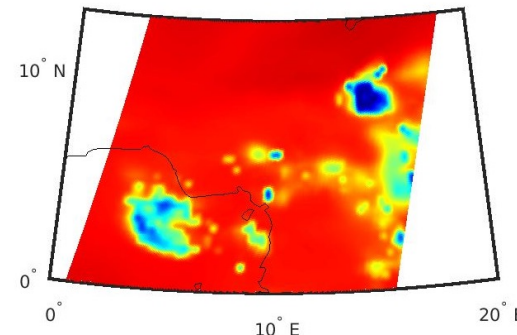
ICI-2V 183.31 GHz
 ± 3.40 GHz



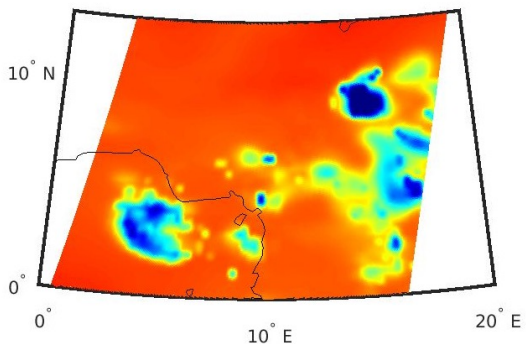
ICI-3V 183.31 GHz
 ± 2.00 GHz



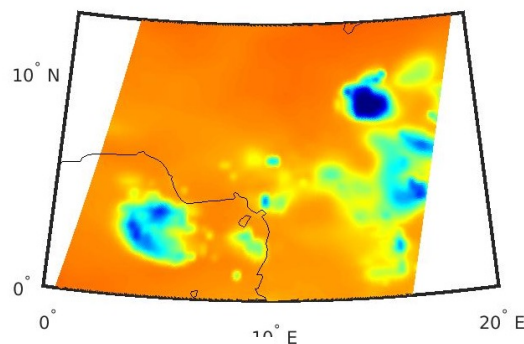
ICI-4H 243.2 GHz
 ± 2.5 GHz



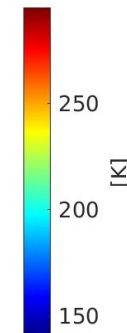
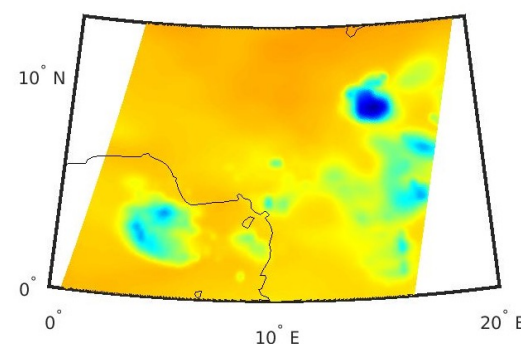
ICI-5V 325.15 GHz
 ± 9.50 GHz



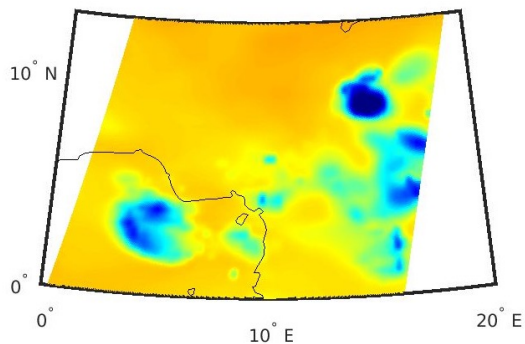
ICI-6V 325.15 GHz
 ± 3.5 GHz



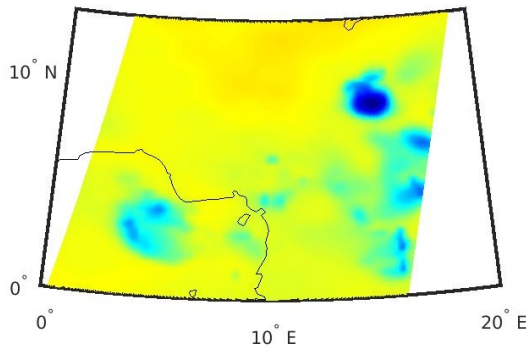
ICI-7V 325.15 GHz
 ± 1.5 GHz



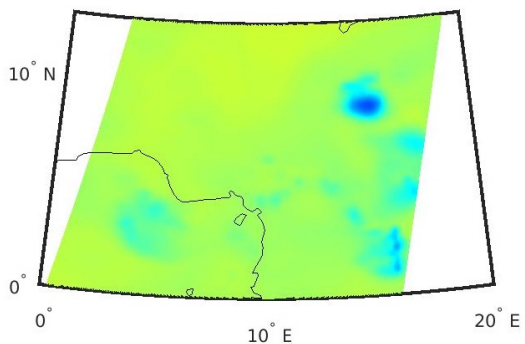
ICI-8V 448 GHz
 ± 7.20 GHz



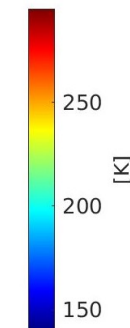
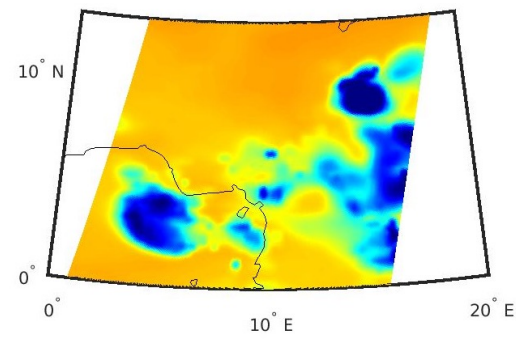
ICI-9V 448 GHz
 ± 3.00 GHz



ICI-10V 448 GHz
 ± 1.40 GHz



ICI-11H 664 GHz
 ± 4.20 GHz





- MWS will expand the MW sounding capabilities in the mid-morning orbit
- ICI provides an unprecedented set of measurements from 183.3 GHz up to 664 GHz, bridging a “spectral gap” between the microwaves and the far infrared. New possibilities!
- MWI (on the same platform with ICI) will continue and enhance important measurements of cloud and precipitation and support surface measurements.
- Synergy among missions, including EPS-SG SCA (and Copernicus CIMR in the long run).
- These missions will:
 - Provide continuation and enhancement of EUMETSAT service from polar orbit.
 - Expand by 20+ years the EUMETSAT contribution to climate data records.
 - Contribute to GPM constellation.

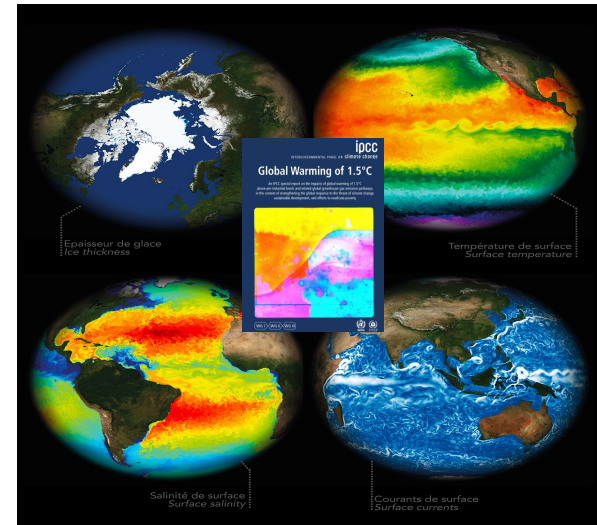
Copernicus Microwave Imager Radiometer (CIMR)



Copernicus Imaging Microwave Radiometer (CIMR)

www.eumetsat.int

- CIMR is a European Space Agency (ESA) High Priority Copernicus mission that responds directly to the *Integrated EU Arctic Policy*
- Conically scanning multi-frequency microwave radiometer in a coordinated flight with MetOp-SG(1B)
- About 95% global coverage every day, mean 6 hourly-revisit in Arctic Areas, dawn-dusk sun-synchronous orbit



MAIN REQUIREMENTS

Channels (GHz, Full Stokes):	1.4	6.9	10.65	18.7	36.5
Footprint (km):	<60	≤15	≤15	≤5.5	<5
NEΔT (K @150K):	≤0.3	≤0.2	≤0.3	≤0.4	≤0.7
Swath	>1900 km				

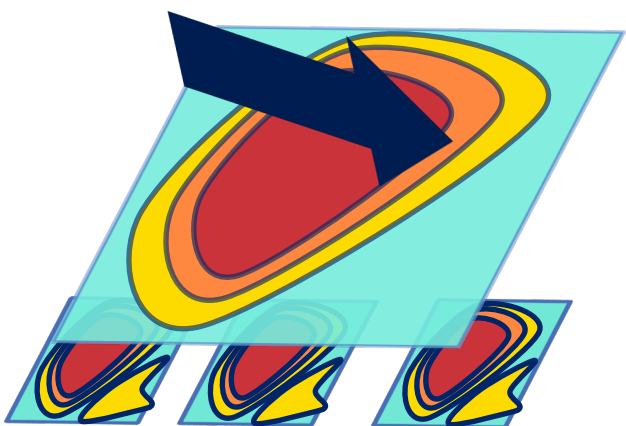
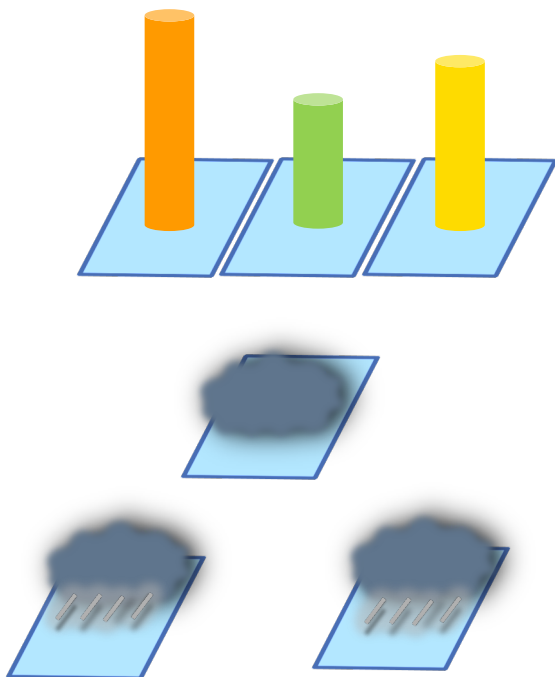




- EUMETSAT will assume responsibility for the production, distribution and archival of **L2 global products over ocean**
- These products are identified in the CIMR Mission Requirements Document (MRD) (see: <https://cimr.eu/documents>)
- L2 global products over ocean will support the activities of Copernicus Marine Environmental Monitoring Service (**CMEMS**) [<http://marine.copernicus.eu>]
Copernicus Climate Change Service (**C3S**) [<https://climate.copernicus.eu/>]



CIMR L2 global products over ocean



Description	Spatial resolution [km]	Total Standard Uncertainty	Timeliness	Revisit
Total Column Water Vapour	<15 km (g=5 km)	<10%	NRT3H	Daily
Liquid Water Path	<15 km (g=5 km)	<50%		
Precipitation rate	<15 km (g=5 km)	<80% at 1 mm h ⁻¹ or <50% above 10 mmh ⁻¹		
Ocean Surface Wind Vector	< 40 km	2 ms ⁻¹ and <20° in direction at wind speeds ≥ 6 ms ⁻¹		
Sea surface Salinity	<60 km	≤0.3 pss		
Sea Surface Temperature	<15 km	≤0.2 K		



- Level-2 Operational Processor for global ocean products to be developed and operated at EUMETSAT
- Providing operational infrastructure for:
 - L1 to L2 data processing;
 - Dissemination;
 - Archival;
 - Performing L2 Cal/Val activities for products over global ocean;
 - Supporting L2 reprocessing.
- Scientific study to develop algorithm prototype in the pipeline, planned KO mid 2023
- Main requirements: physical consistency, multi-frequency, multi-scale approach



- The CIMR orbit selection allows for the possibility of direct synergy at high latitudes with EPS-SG and other Copernicus missions
- This needs to be explored to enhance the impact of EPS-SG and Copernicus data
- Elsewhere, complementarity is the word: global NWP/Ocean models taking benefit from assimilation of EPS-SG and Copernicus data
- Higher level products (Level 3, Level 4)
- Benefit of combined retrievals, inter-comparison/validation of geophysical parameters

EPS Sterna



- Since December 2018 EUMETSAT identified small microwave sounding satellites complementing the reference EUMETSAT Polar System (EPS–SG) /Microwave Sounder (MWS) mission as a possible additional EUMETSAT contribution to the realisation of the WIGOS Vision 2040, in line with objective 4 of the EUMETSAT strategy “Destination 2030”.
- In 2019, EUMETSAT prioritised relevant user requirements, starting from the agreed EPS–SG/MWS user requirements.
- The European Space Agency (ESA) Arctic Weather Satellite (AWS) programme was approved at the ESA ministerial Council in 2019.
- See: https://www.esa.int/Applications/Observing_the_Earth/Meteorological_missions/Arctic_Weather_Satellite
- The AWS observational requirements are in line with EUMETSAT requirements’ prioritization and were endorsed by the EPS–SG MWS/AWS Science Advisory Group.



- The AWS is a small microwave sounding satellite planned to be launched in Q1/Q2 2024, with the objective to:
 - Provide a one year in-orbit demonstration for a future, operational constellation;
 - Confirm the impact of increased passive microwave soundings on Numerical Weather Prediction (NWP) accuracy and Nowcasting, particularly over the Arctic;
 - Demonstrate a cost-effective approach.
- If successful, the AWS prototype (PFM) will constitute an opportunity for EUMETSAT to implement a constellation of small sounding microwave satellites (i.e. EPS-Sterna) in complementary orbits to EPS-SG and expand the products' envelope of the EPS-SG mission for its users .
- The EPS-Sterna constellation will be developed in cooperation with ESA
- Composed of recurrent small satellites stemming from the AWS PFM
- Programme approval targeted in 2025 after one year in orbit demonstration.



AWS payload channels and performance requirements

Channel	Frequency (GHz)	Bandwidth (MHz)	NEΔT (K)	Footprint (km)	Utilisation
AWS-11	50.3	180	<0.6	≤ 40 km	Temperature sounding
AWS-12	52.8	400	<0.4	≤ 40 km	Temperature sounding
AWS-13	53.246	300	<0.4	≤ 40 km	Temperature sounding
AWS-14	53.596	370	<0.4	≤ 40 km	Temperature sounding
AWS-15	54.4	400	<0.4	≤ 40 km	Temperature sounding
AWS-16	54.94	400	<0.4	≤ 40 km	Temperature sounding
AWS-17	55.5	330	<0.5	≤ 40 km	Temperature sounding
AWS-18	57.290344	330	<0.6	≤ 40 km	Temperature sounding
AWS-21	89	4000	<0.3	≤ 20 km	Window and Cloud detection
AWS-31	165.5	2800	<0.6	≤ 10 km	Window/humidity sounding
AWS-32	176.311	2000	<0.7	≤ 10 km	Humidity sounding
AWS-33	178.811	2000	<0.7	≤ 10 km	Humidity sounding
AWS-34	180.311	1000	<1	≤ 10 km	Humidity sounding
AWS-35	181.511	1000	<1	≤ 10 km	Humidity sounding
AWS-36	182.311	500	<1.3	≤ 10 km	Humidity sounding
AWS-41	325.15±1.2	800	<1.7	≤ 10 km	Humidity sounding/cloud detection
AWS-42	325.15±2.4	1200	<1.4	≤ 10 km	Humidity sounding/cloud detection
AWS-43	325.15±4.1	1800	<1.2	≤ 10 km	Humidity sounding/cloud detection
AWS-44	325.15±6.6	2800	<1	≤ 10 km	Humidity sounding/cloud detection

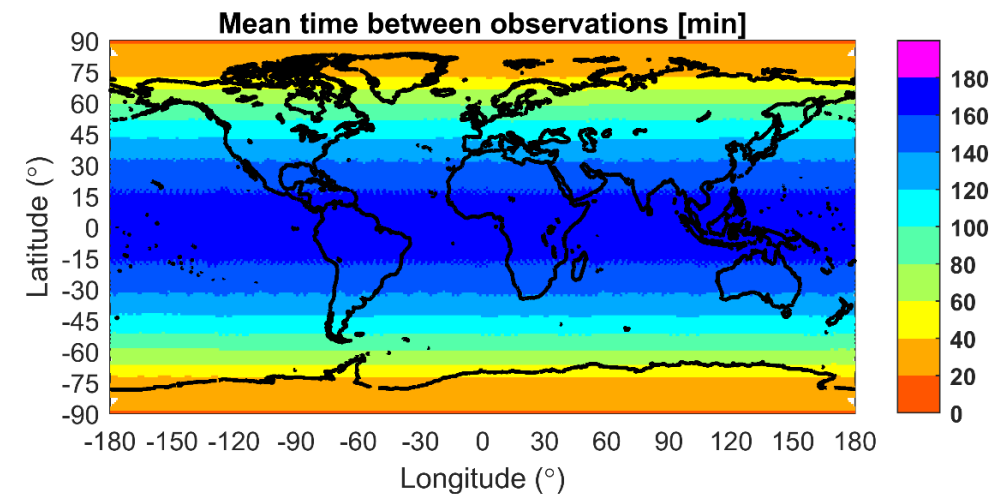
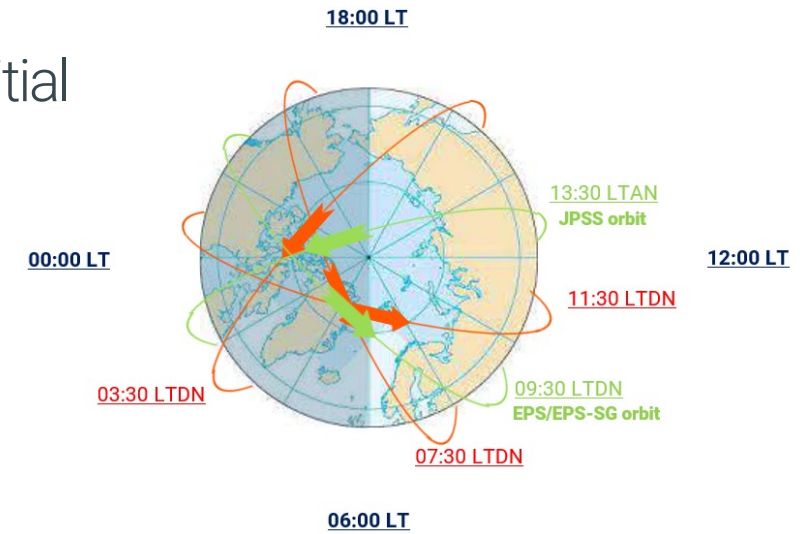
NEΔT comparable to legacy missions (e.g. EPS-SG MWS/JPSS ATMS)

Radiometric accuracy shall be <1K

Lifetime stability < 0.2 K



- Constellation architecture
 - 6 satellites in 3 orbital planes (2 satellite in each plane) for initial constellation
 - Planes chosen to be complementary to EPS/EPS-SG and JPSS orbits (LTDN: 03:30, 07:30, 11:30)
- Constellation performance
 - Time to achieve 90% global coverage (driving requirement):
 - 3.10 - 4.7 hours- Sterna constellation alone
 - 2.37 - 3.5 hours- Sterna +EPS-SG and JPSS
 - Mean time between observations: 20 minutes - 3 hours





- The EPS Sterna Constellation will focus on:
 - Water vapor sounding
 - Temperature sounding
 - Cloud & precipitation detection
- The primary objectives of the EPS Sterna Constellation would be to:
 - Contribute to improved global and regional NWP accuracy
 - Complement the microwave observations from the Metop/Metop-SG and NOAA JPSS (Joint Polar Satellite System) polar-orbiting meteorological satellites
- The mission will also contribute to nowcasting applications over the Arctic region through the increase in the frequency and availability of microwave observations.
- Climate monitoring is not a primary objective of the mission, however EPS Sterna is expected to contribute through the provision of long-term data records and implementation of the necessary capabilities.



Thank you!
Questions are welcome.